

Daytime seasonal variation of Longitudinal structures of electron density and temperature measured with DEMETER satellite

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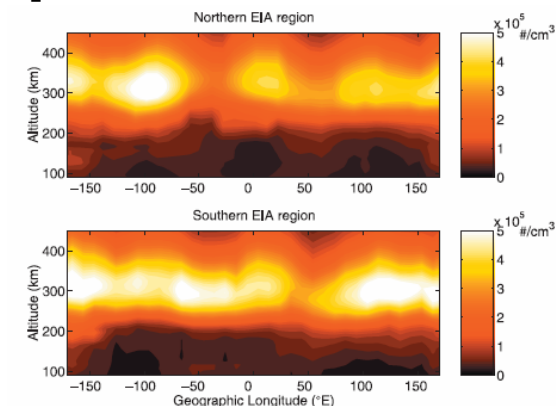
Content

- Introduction
 - Recent observations of wave-4 longitudinal structure
 - Motivation
- Longitudinal structure of Ne & Te observed by **DEMETER**
 - Focus on **seasonal** variation under **low solar activity**
- **Longitudinal structure of Ne & Te observed by Hinotori**
 - Focus on **LT** variation under **high solar activity**
- Summary

Introduction :observations of longitudinal structure

- Air glow [*Sagawa et al.*, 2005; *Immel et al.*, 2006]
- Equatorial electro jet [*England et al.*, 2006]
- Total electron content [*Scherliess et al.*, 2008]
- Plasma drift [*Kil et al.*, 2007]
- Neutral wind at lower thermosphere [*Forbes et al.*, 2003]
- Ne profile obtained by radio occultation [*Lin et al.*, 2007]
- In-situ plasma observaton [*Liu & Watanabe* 2008]
- Neutral density [*Liu et al.*, 2009]
- Ne in the top side ionosphere [Bankov et al., 2009]
- Etc...

Many observations have confirmed the wave-4 longitudinal structure



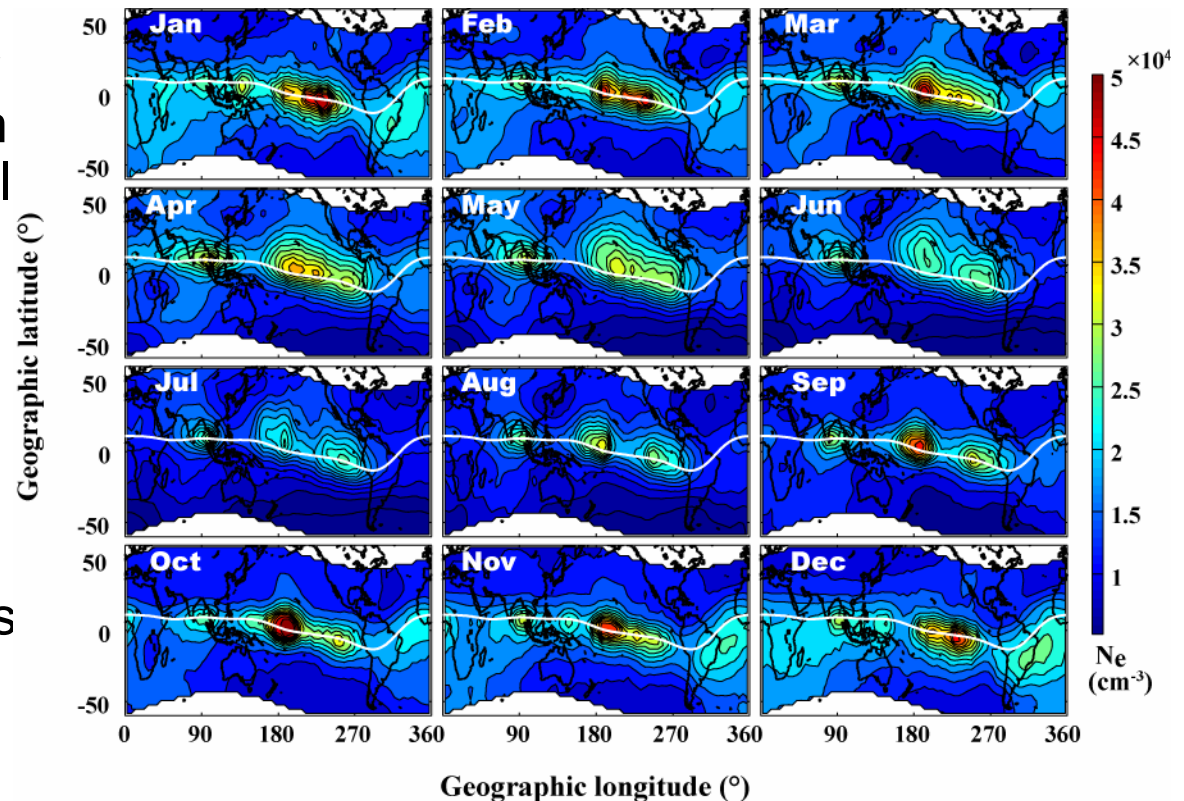
[*Lin et al.*, 2007]

Motivation

- Longitudinal structures of Te & Ne in the topside ionosphere have been reported by *Ren et al.* [2008]
 - However, ...
 - Using 11 years integrated data of DMSP(800km)
 - Data were divided into 4 season
 - They concluded that Te variation is related to in-situ Ne variation
- Therefore, these results might be included solar activity variation
- Further, we confirm detail of seasonal variation of longitudinal structure of Ne & Te
- We found the evidence Te in the topside ionosphere does not necessary correlate with in-situ Ne

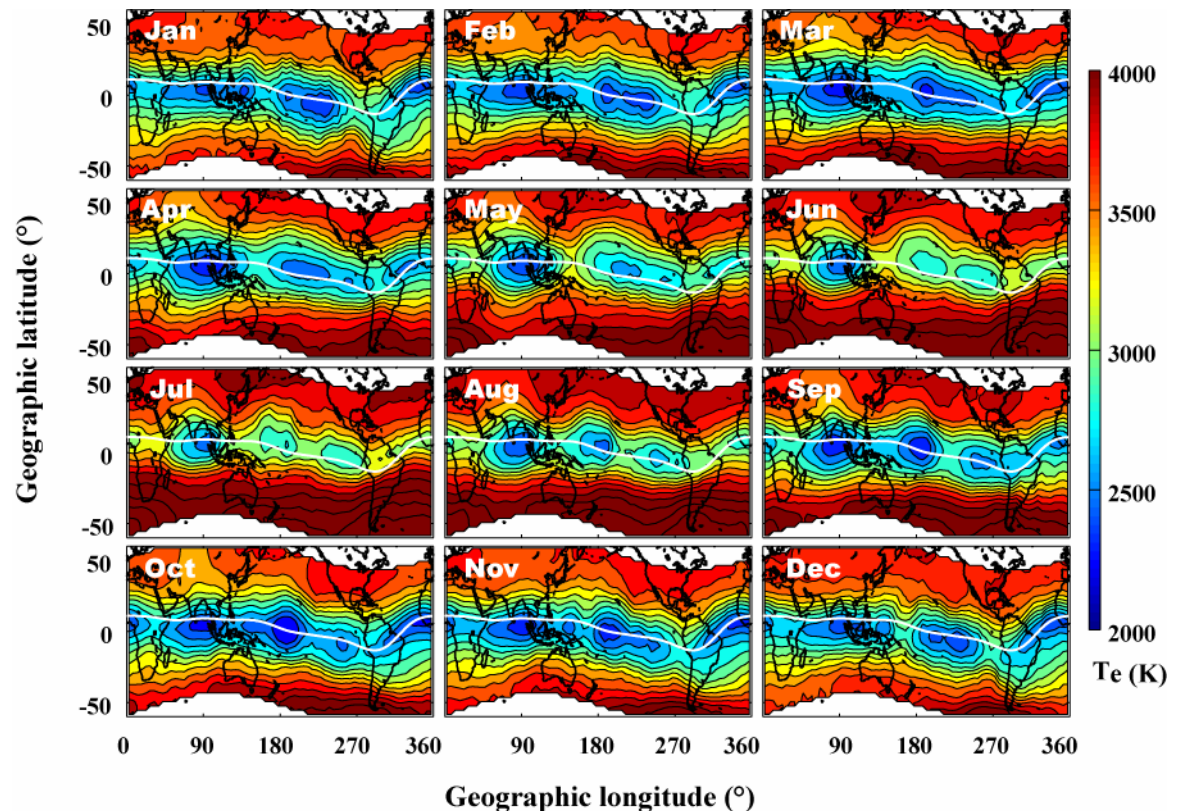
Seasonal variation observed by DEMETER \sim Ne

- Maximum Ne is high around magnetic equator in December sol. & low in June sol. due to seasonal variation of O/N_2
- Ne in summer hemisphere is higher than that in winter hemisphere due to neutral wind
- Longitudinal structure has seasonal variation
 - Wave-4: May-Oct.
 - Wave-3: Nov.-Apr



Seasonal variation observed by DEMETER $\sim T_e$

- T_e generally shows negative correlation of T_e
- Minima of T_e are high around magnetic equator during Jul.-Aug. & low during Oct.-May due to Ne variation
- However, it does not correspond exactly
- Longitudinal variation
 - Wave-4: Jul.-Sep.
 - Wave-3: Oct.-Jun.



Ne & Te in the magnetic equator (|dip Lat.|<5)

- maximum of Ne are seen around 90, 190, 250 and 330 °E
 - 190°E : high during Aug.–Mar.
 - 330°E : low during Sep.–Apr. and almost disappears during May–Aug.
- the minima of Te are seen around 90, 190, 250, and 330°E.
 - 90°E: not always the lowest
 - 190°E: Feb.–Sep.
- Although the maximum of Ne at 190°E is larger than the other maxima of Ne during
- Te variation does not correspond to Ne exactly

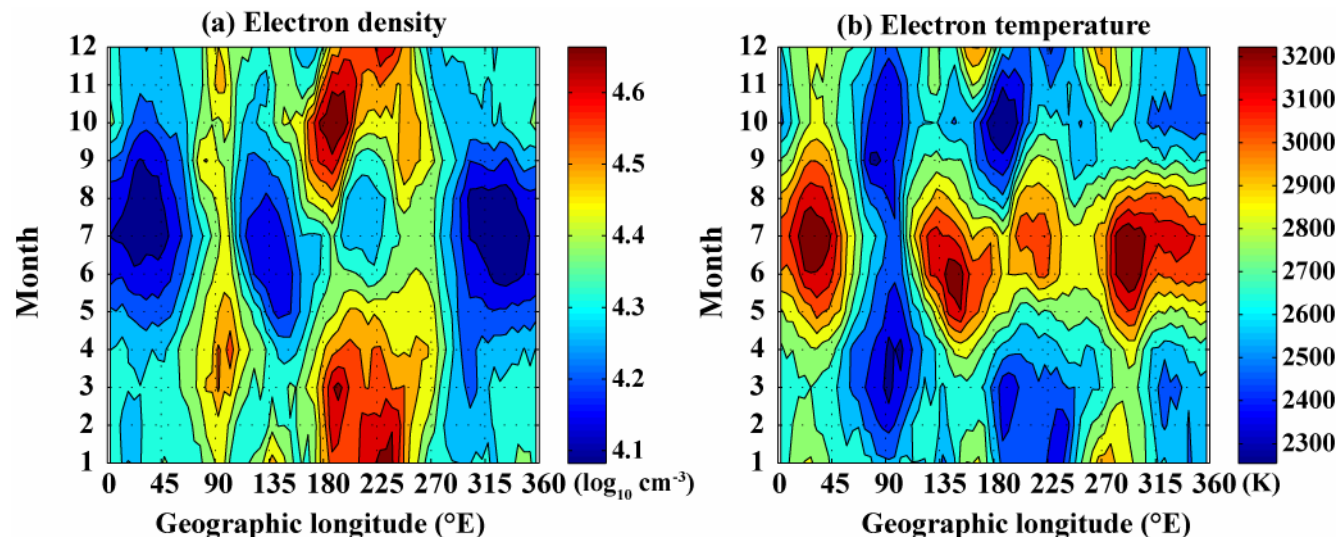
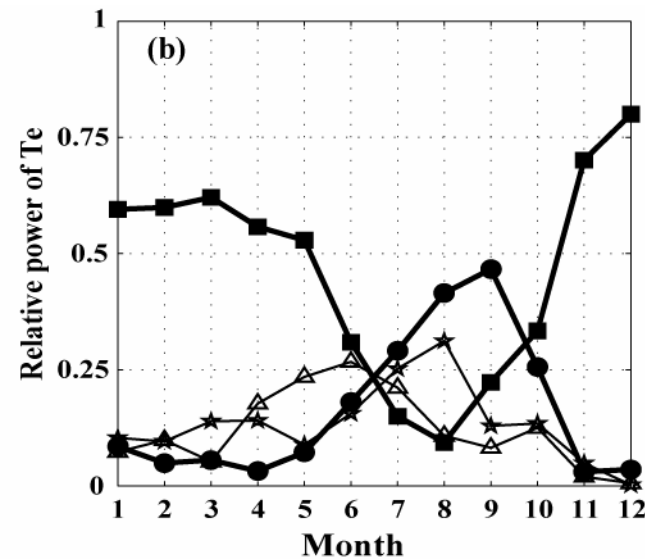


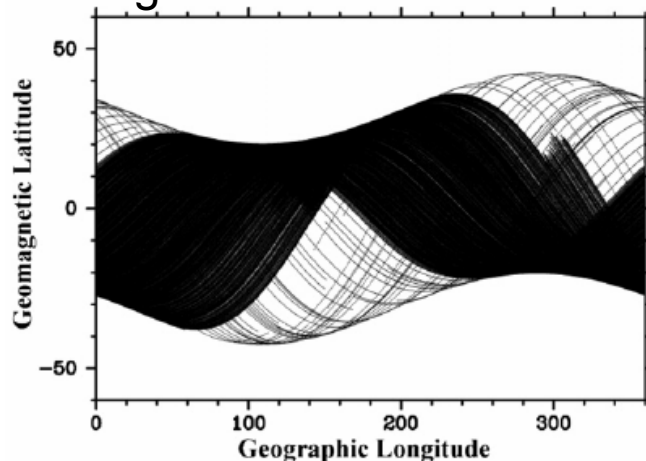
Fig.8



HINOTORI

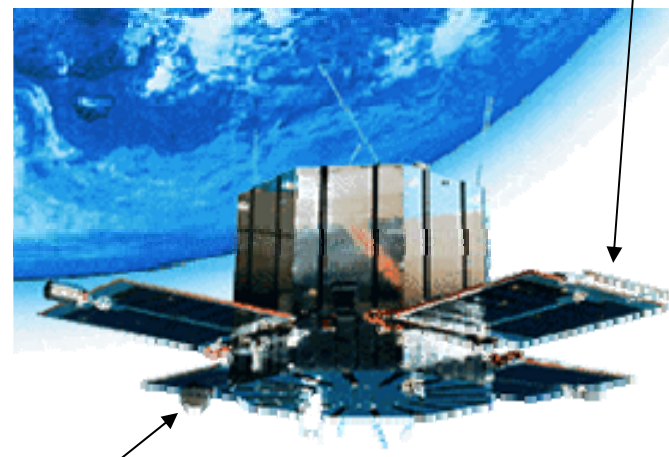
- Japanese satellite
- 23 Feb.1981-14 June 1982
- Solar activity :maximum and medium
- Latitude: -32° ~ 32°
- Height: 576~644 km
- Electron Density (Ne) & Temperature (Te)
- All LT range

Fig. 2



[Kakinami et al., 2008]

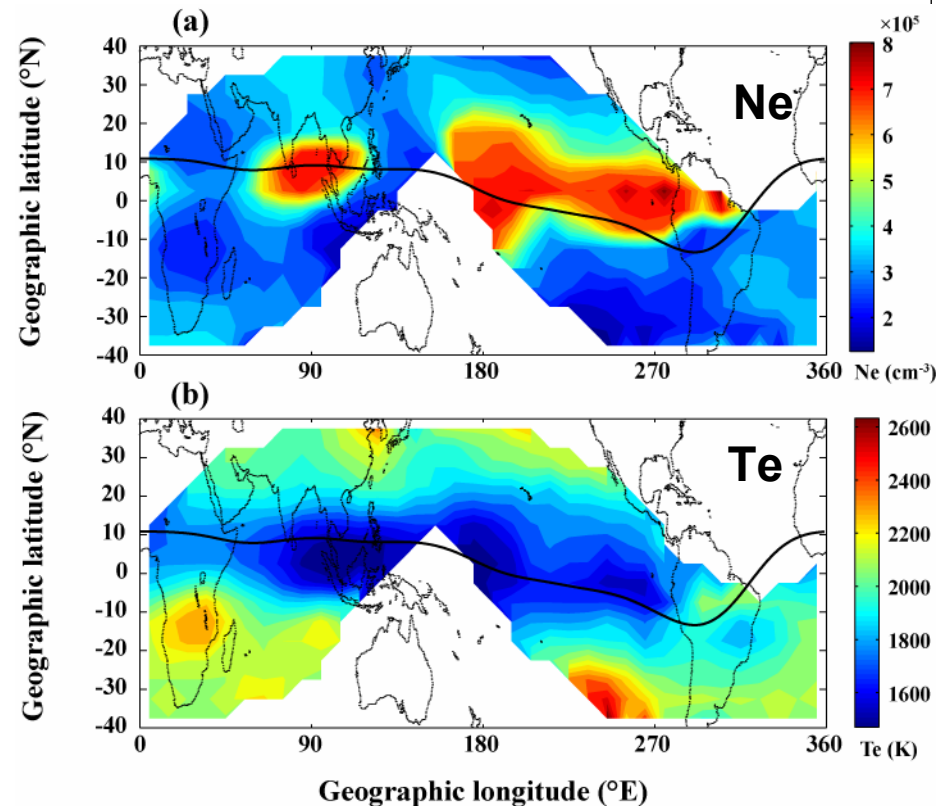
Impedance Probe [Oya, et al., 1986]



Electron temperature Probe [Hirao and Oyama, 1970]

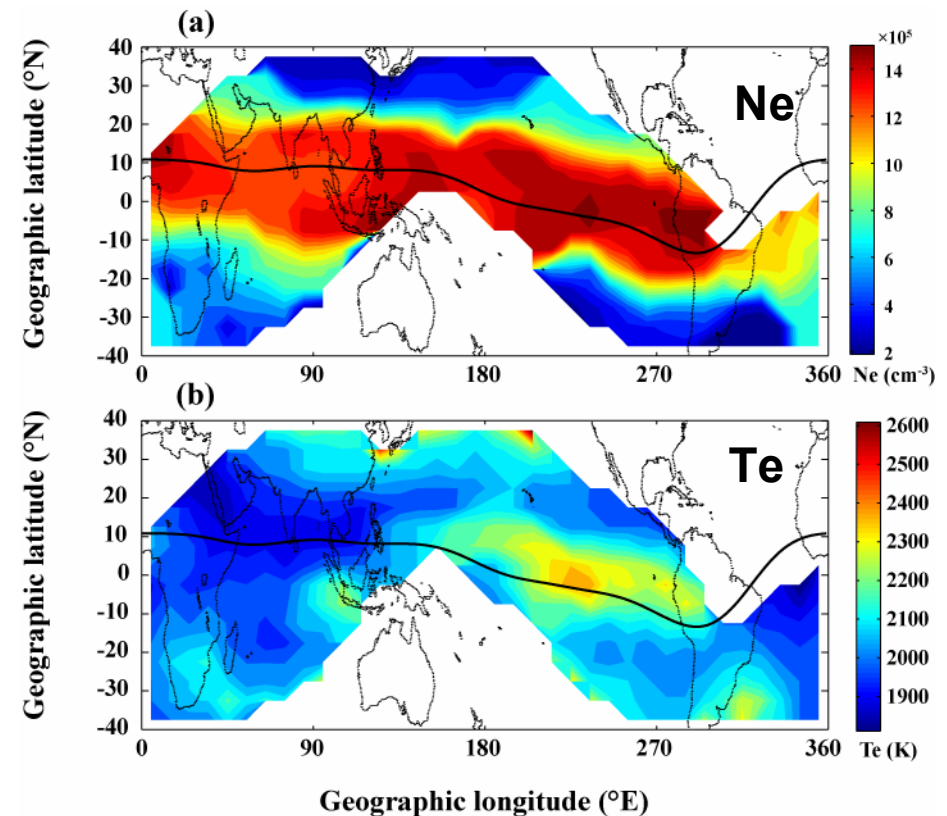
LT variation observed by HINOTORI (1)

- Median during **July-October** under $K_p < 4$
- LT=0900-1100
- Wave-4 like structure can be seen
- Near the magnetic equator, correlation between Ne & Te is negative



HINOTORI ~LT variation observed by HINOTORI(2)

- The same condition of Fig. 4 but different LT
- LT=1300-1500
- 4 maxima of Ne are seen
- Peak Ne are shifted to eastward
 - Basic feature of nonmigrating tide of wave number 4 (DE3)
- In the magnetic equator, correlation between Ne & Te is positive



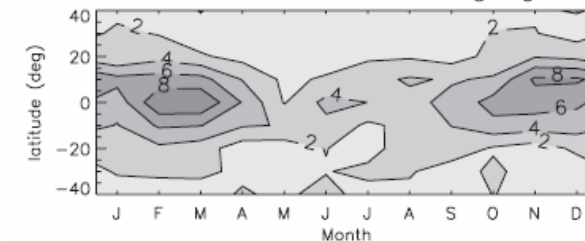
Discussion

- Seasonal variation of wave-3/-4 structure of Ne & Te is similar to DE2/DE3 zonal wind at dynamo region
- However, seasonal variations of Ne & Te show discrepancy
- **Cause of Ne variation**
 - Electric field which produces the fountain effect is modulated by zonal wind [*Jin et al.*, 2008]
 - Ne in the topside is also affected by meridional wind in the topside ionosphere [*Watanabe & Oyama*, 1996]
- **Cause of Te variation**
 - Heat conduction is dominant in the top side ionosphere
 - Te is correlated with integrated Ne along the magnetic field line below the satellite height in the topside [*Kakinami et al.*, 2011 submitted to JGR]
 - Therefore, lower ionosphere condition might reflect Te in the topside ionosphere
- Meridional wind might produce the discrepancy

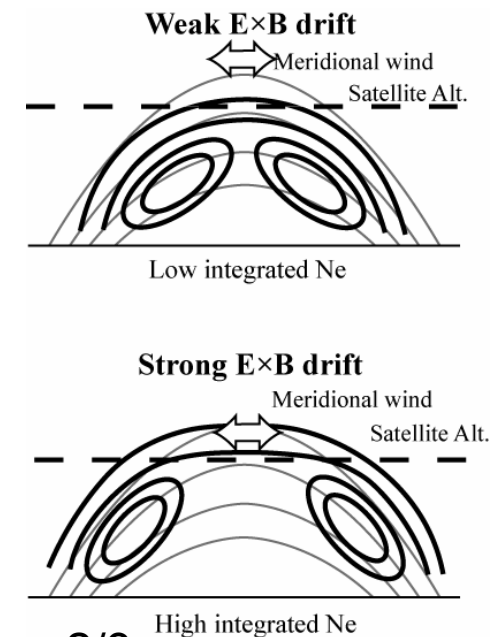
Zonal wind DE3@95km



Meridional wind S=-3



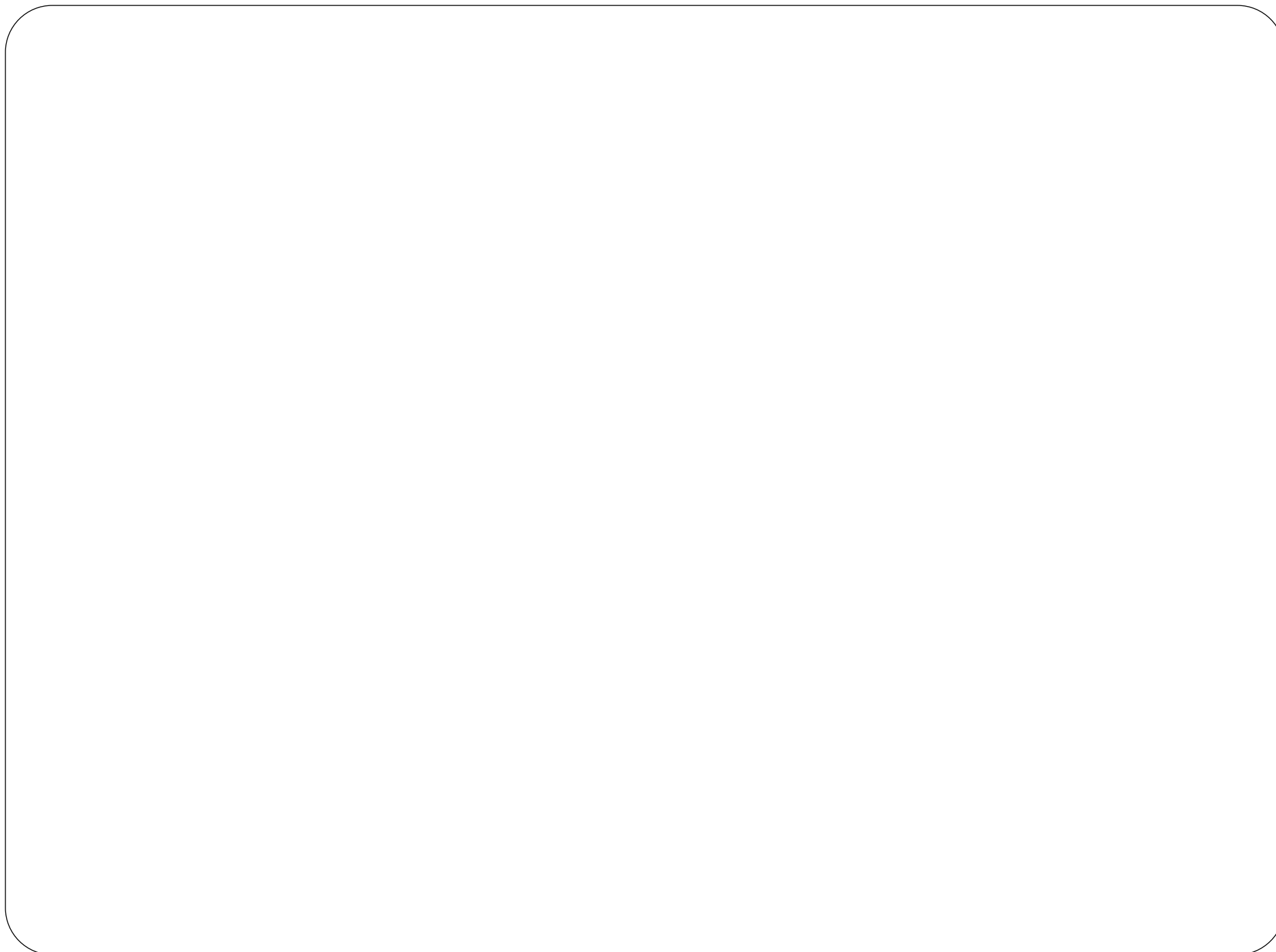
[*Forbes et al.*, 2003]



DE2/3: eastward propagating diurnal tide with wave number 2/3

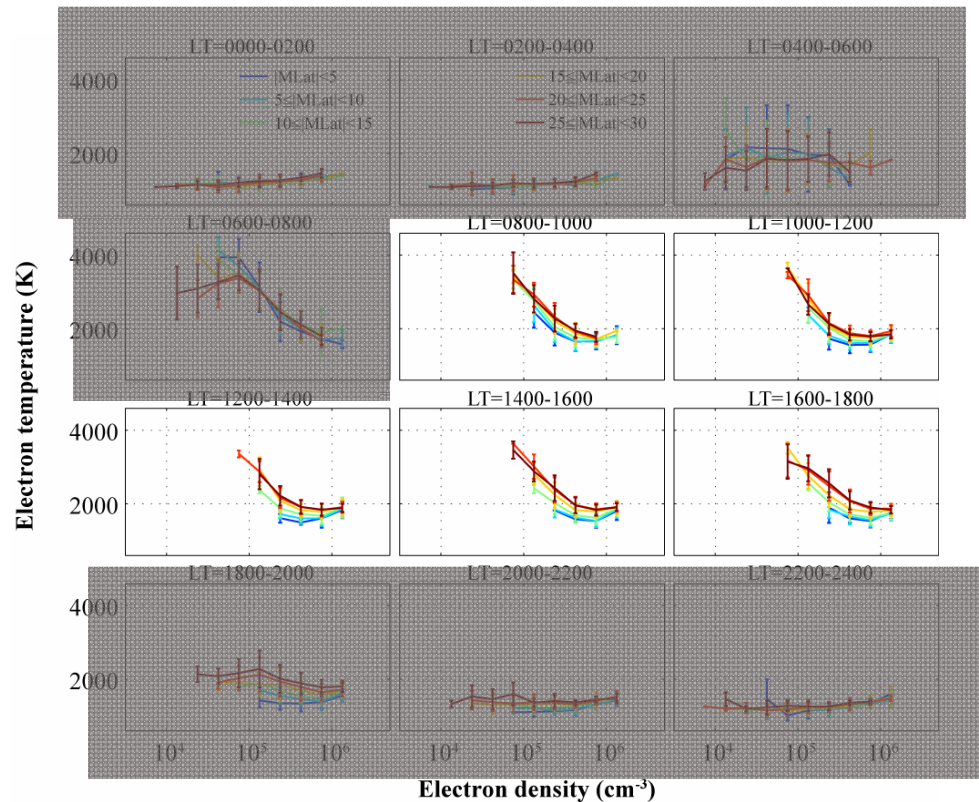
Summary

- Longitudinal structure (wave-3/-4) are seen during both low and high solar activity
- Wave-3/-4 of both Ne & Te are seen when zonal component of DE2/DE3 is enhanced
- One of 4 or 3 Ne peaks are pronounced
 - Only Strong fountain effect which can reach the satellite height can be detected
- Discrepancy between Ne & Te
 - Ne □ May-Oct. (wave-4) □ Nov.-Apr. (wave-3)
 - Te □ Jul.-Sep. (wave-4) □ Oct.-Jun. (wave-3)
 - Maximum of Ne often appear around 190 deg.
 - Minimum of Te often appear around 90 deg.
 - Therefore, cause of longitudinal structure is different between Ne & Te
- In the topside ionosphere, correlation between Ne & Te is not always negative
 - The correlation is positive when Ne is significantly high
- Ne easily affected by meridional wind
- Therefore, Te is a better indicator for nonmigrating tidal effect at topside ionosphere than the Ne.

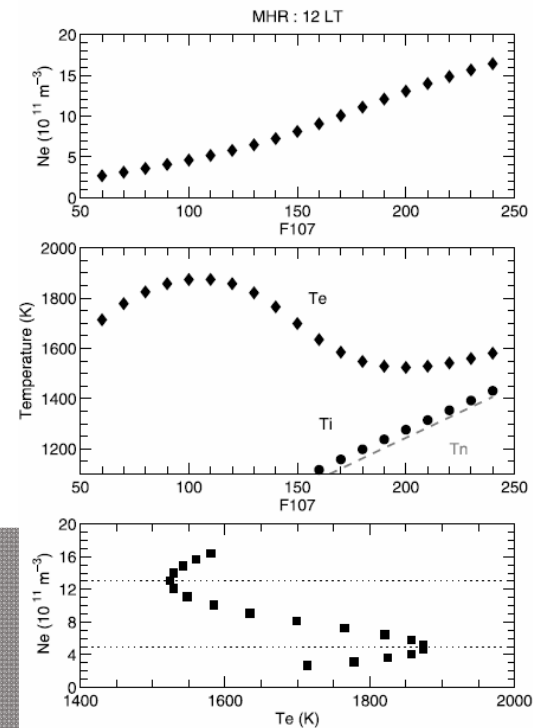


Discussion (2)

- Positive correlation between Ne & Te
 - No significant seasonal variation
 - Magnetic latitude dependency



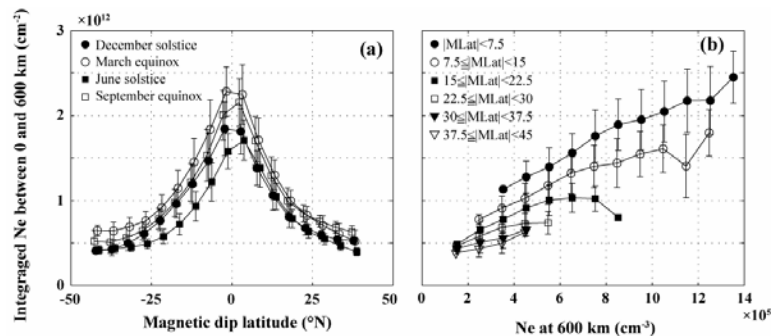
Hinotori@600km
each line indicate magnetic latitude region



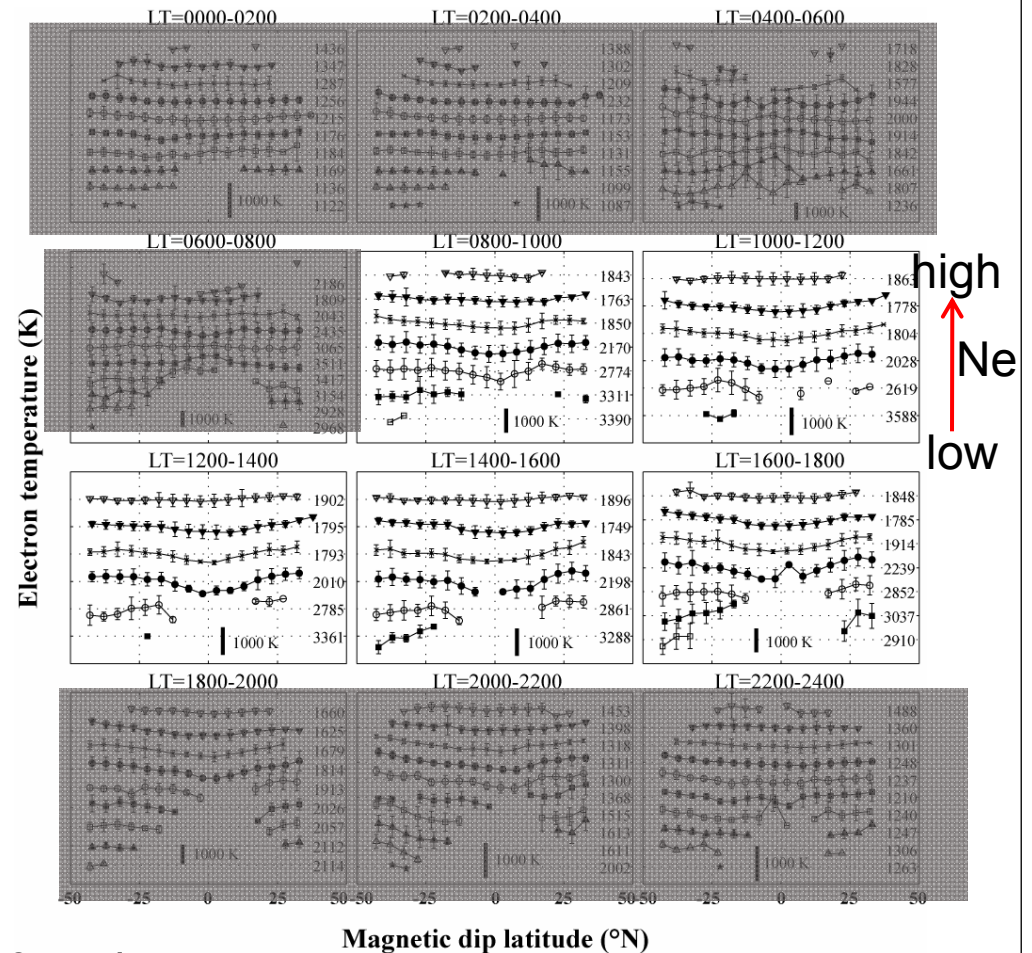
Simulation@F2 peak
[Lei et al., 2007]

Discussion (3)

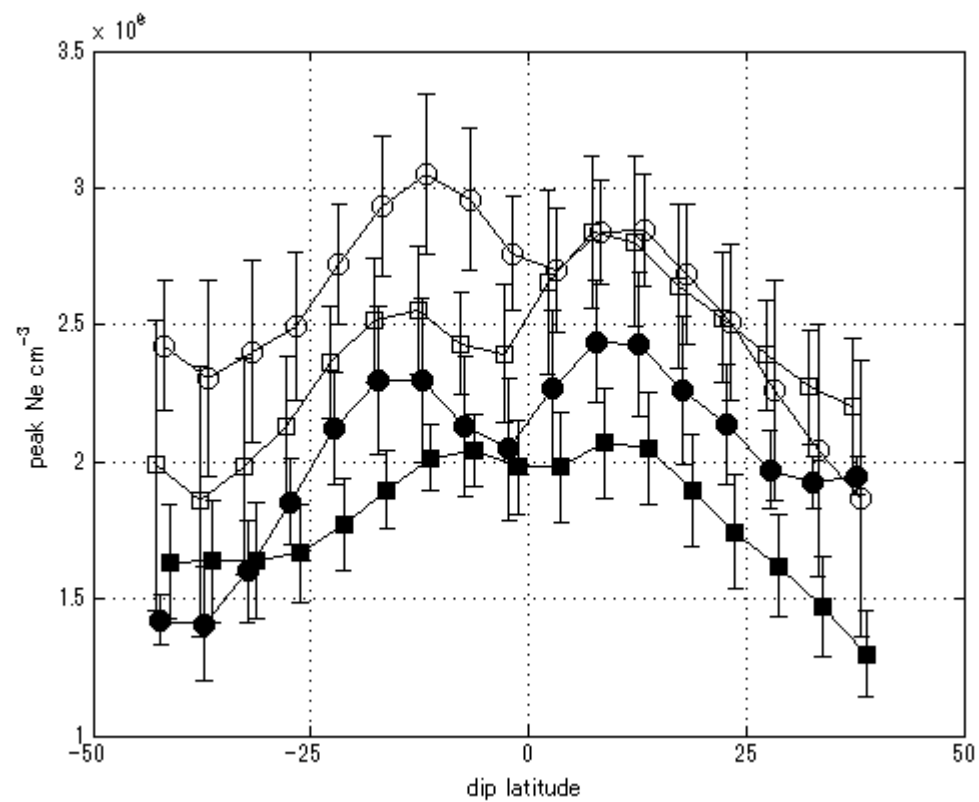
- Each symbol indicate same Ne region
- Te decreases with increasing magnetic Lat. during daytime when Ne is the same region
 - Te does not fully depend on in-situ Ne rather integrated Ne



Integrated Ne along field line between 0 and 600 km below the satellite height



Discussion (2)



● December solstice, ○ March equinox
■ June solstice, □ September equinox

DEMETER

Fig. 3



- French satellite
- Detection of Electro-Magnetic Emissions Transmitted from Earthquake Regions
- January 2004 □
- Solar activity: low
- Magnetic field, Electric field (DC to 3.25MHz), T_e & ion temperature (T_i), N_e & ion density (N_i), ion velocity, electron spectrum (> 30 KeV)
- Inclination: 98° , sun-synchronous (1030, 2230 LT)
- Altitude: ~ 710 km, 660 km after December 2005
- In this study, data in 2006 & 2007 are used